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THESIS

THE EFFECTS OF LIQUID PROPELLANT MOTION ON THE ATTITUDE STABILITY OF SPIN STABILIZED SPACECRAFT

bу

Jack W. Myers, Jr.

March 1990

Thesis Advisor:

Brij N. Agrawal

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The Effects of Liquid Propellant Motion on the Attitude Stability of Spin Stabilized Spacecraft

by

Jack W. Myers, Jr. Major, United States Army B.S., United States Military Academy

Submitted in partial fulfillment of the requirements for the degree of

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B. N. Agrawa

Author:

Jack W. Myers, Jr.

Approved by:

Brij N. Agrawal, Thesis Advisor

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Michael, R. Gorman, Second Reader

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Department of Aeronautics and Astronautics

ABSTRACT

An analysis of the effects of liquid motion on the attitude stability of spin stabilized spacecraft is presented. The effects of varying the fuel load and the asymmetry of the platform are emphasized. The energy sink stability criteria are derived and applied to a marginally stable spacecraft. The stability predictions based on the energy sink stability criteria are compared to the results of a computer simulation. Based on this comparison the limitations of the energy sink stability criteria are identified.



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I. INTRODUCTION

A. BACKGROUND

Historically, the attitude stability conditions of spinning spacecraft have been derived using the energy sink method. A significant shortcoming of this method is that it does not account for the dynamic interaction of liquid motion. For current spacecraft with large amounts of liquid propellant, dynamic interaction can destabilize the spacecraft. As a result, under some circumstances the stability prediction of the energy sink method can be in error.

Despite this shortcoming, the energy sink method continues to be in wide spread use as an analytical technique for determining attitude stability for dual spin spacecraft. This being the case, it is essential that the nature of the deficiency and the conditions under which it occurs be understood.

B. OBJECTIVE

The objective of this study is to investigate and more accurately define the nature of this shortcoming. Specifically, the case of a marginally stable, dual spin spacecraft with a despun platform will be explored. The

stability of the spacecraft will be examined as the percent of platform asymmetry and the fuel load are varied. Marginally stable, in terms of the energy sink stability conditions, means the spacecraft has an inertia ratio slightly greater than one. The percent asymmetry is defined as the ratio of the difference of the two transverse moments of inertia to the sum of the two transverse moments of inertia times 100. The fuel load is the amount of fuel on board expressed as a fraction of the total fuel capacity of the spacecraft.

To perform the analysis, a computer simulation developed by Chung [Ref. 1] of a dual spin spacecraft will be used to determine the stability of the various configurations. These results will be compared to the predictions of the energy sink stability criteria as developed by Likins in Reference 2.

C. LITERATURE REVIEW

The energy sink approach was first applied to study the effects of energy dissipation on the stability of a freely spinning body in 1963 by Thomson and Reiter [Ref. 3]. This led to the well known requirement for stability, an object must spin about its axis with the largest principal moment of inertia.

In 1966, Likins [Ref. 2] developed the energy sink stability conditions for asymmetric, dual spin spacecraft

assuming the energy transfer between the platform and the rotor was small. His results showed that the average energy dissipation on the platform or on the rotor is proportional to the average of the rate of change of the corresponding inertial spin rate. He went on to point out that the ratio of the spin moment of inertia to the algebraic mean transverse moment of inertia was the critical stability quantity.

In 1972, two papers questioned Likins' conclusions suggesting that the key stability parameter was the ratio of the spin moment of inertia to the geometric mean transverse moment of inertia rather than the algebraic mean transverse moment of inertia. In his 1974 paper [Ref. 4]. Spencer's analysis corroborated the importance of the geometric mean transverse moment of inertia.

In 1981. Hubert [Ref. 5] concluded that using core energy instead of total energy in the expression of energy dissipation, made the energy sink prediction applicable to a dual spin satellite with energy dissipating devices on the platform.

In 1983, Cochran and Shu [Ref. 6] used the generalized method of averaging to study both energy dissipation and the energy addition required to maintain the constant spin rate of the rotor. Their conclusions substantiated Hubert's

hypothesis concerning core energy assuming that the internal mass motion is sufficiently small.

In Reference 1, Chung studied the application of the energy sink method to the INTELSAT VI satellite. His conclusion was that the energy sink method did not correctly predict stability for all cases. The results of Reference 7, also a study of the INTELSAT VI satellite, indicated that stability increased as the fuel load increased, and that stability decreased as the platform asymmetry increased.

The energy sink method continues to be used throughout the community to determine the stability of spinning spacecraft. In order to have confidence in the predictions of the method, it is necessary to understand its limitations. It is hoped that this study will provide additional insight into the energy sink method's transition zone. In other words, where the boundary is between accurate energy sink predictions and erroneous predictions.

D. ORGANIZATION OF STUDY

Chapter II presents the derivation of the energy sink attitude stability criteria for an asymmetric dual spin spacecraft, a description of the simulation used to model the satellite motion, and a description of the satellite configuration used for the study. Chapter III describes the development of the system parameters for the spacecraft

configuration under study, the application of the energy sink stability criteria to determine a stability prediction, and the methodology used to determine what conditions to pirclate. Chapter IV presents the results and analysis of both the simulation and the energy sink predictions. Chapter V summarizes the conclusions based on the results presented in Chapter IV.

II. BACKGROUND

The analysis conducted in this study involves the comparison of the energy sink stability predictions with the results of a computer simulation. Both require a satellite configuration in order to define their input. This chapter covers the derivation of the energy sink stability criteria, a description of the simulation used for the study and a description of the satellite configuration.

A. ENERGY SINK DERIVATION

The general Likins' model [Ref. 2] used to derive the stability conditions consists of an asymmetric body P and an axisymmetric body R (Figure 1).



Figure 1. Idealized dual spin system.

With the coordinate system located at the center of mass of the total system, and fixed in the body P such that the center of mass of both bodies lies along the \hat{a}_3 axis, the rotational equation of motion governing the system is.

$$\mathbf{M} = d\mathbf{H}/d\mathbf{t} = 0 \tag{1}$$

where M is the total external moment exerted on the system and H is the total angular momentum of the system. The resulting linearized equations of motion are,

$$M_{1} = I_{1} \dot{w}_{1} + (I_{p} - I_{0}) w_{p} w_{0} + I_{0} w_{p} w_{0} = 0$$
 (2)

$$M_{2} = I_{2}\dot{w}_{2} + (I_{1} - I_{2})w_{2}w_{1} - I_{2}w_{3}w_{3} = 0$$
(3)

$$M_{0} = I_{1}w_{0} + I_{1}w_{1} + I_{1}w_{1}w_{0} - I_{1}w_{1}w_{0} = 0 \qquad (4)$$

where I_{ij} is the total principal axial moment of inertia of the system, I_{ij} is the principal axial moment of inertia of the body R. I_{ij} and I_{ij} are the principal transverse moments of inertia of the system, w_{ij} is the axial component of the angular velocity of the body P, w_{ij} is the relative rate at which the body R is moving with respect to the body P, and w_{ij} and w_{ij} are the transverse components of the total system's angular velocity. These three equations have four unknowns: w_{ij} , w_{ij} , w_{jj} , and w_{jj} . The fourth equation describes the rotor/platform interface,

$$\mathbf{T}_{\mathrm{m}} = \mathbf{I}_{\mathrm{r}}(\dot{\mathbf{w}}_{\mathrm{p}} + \dot{\mathbf{w}}_{\mathrm{r}}) \tag{5}$$

where T_m is the resultant moment of all the forces acting by P on R about the rotor axis. Equations (2) and (3) can be

rewritten as,

$$\dot{w}_1 + \sigma_1 w_2 = 0 \tag{(2)}$$

$$\dot{w}_{1} - \sigma_{1}w_{2} = 0 \qquad (7)$$

where,

$$\sigma_{1} = [(I_{p} - I_{1})w_{p} + I_{p}w_{p}]/I_{1}$$
(3)

$$\sigma_{1} = [(I_{1} - I_{1})w_{1} + I_{1}w_{1}]/I_{1}$$
(9)

Letting $w_1(0) = w_1$ and $w_1(0) = 0$ yields the following solution to equations (E) and (7),

$$w_1 = -w_2 \sqrt{(\sigma_1/\sigma_2)} \sin \sqrt{(\sigma_1\sigma_2)} t$$
 (10)

$$w_{\bullet} = w_{\bullet} \cos((\tau \cdot \sigma_{\bullet})t)$$
 (11)

Applying the Routh-Hurwitz stability criterion to the characteristic equation of equations (6) and (7) shows that for stability,

$$\sigma \cdot \sigma - > 0 \tag{12}$$

The equations for the angular momentum and rotational kinetic energy for the system described above are,

$$h^{2} = I_{1}^{2} w_{1}^{2} + I_{2}^{2} w_{2}^{2} + (I_{p} w_{p} + I_{r} w_{r})^{2}$$
(13)

$$2T = I_1 w_1^2 + I_2 w_2^2 + I_p w_p^2 + I_r w_r^2 + 2I_r w_r w_p$$
(14)

The nominal angular momentum is given by,

$$h_{j} = I_{p}w_{p} + I_{r}w_{r} \qquad (15)$$

and is assumed to be constant. Taking the derivative of equations (13) and (14) with respect to time, and taking into consideration conservation of angular momentum and assuming an energy dissipation mechanism yields,

$$0 = I_{1}^{2} w_{1} \dot{w}_{1} + I_{2}^{2} w_{1} \dot{w}_{2} + (I_{p} w_{p} + I_{p} w_{2}) (I_{p} \dot{w}_{p} + I_{p} \dot{w}_{2})$$

$$\dot{T} = I_{1} w_{1} \dot{w}_{1} + I_{2} w_{2} \dot{w}_{2} + I_{p} w_{p} \dot{w}_{p} + I_{v} w_{v} \dot{w}_{v} + I_{v} \dot{w}_{v} \dot{w}_{v} \dot{w}_{v} + I_{v} \dot{w}_{v} \dot{w} \dot{w}_{v} \dot{w}_{v} \dot{w} \dot{w}_{v} \dot{w} \dot{w}_{v} \dot{w} \dot{w}_{v} \dot{w} \dot{w}_{v} \dot{w} \dot{w}_{$$

Now the solutions in equations (10) and (11) are no longer correct. However, if the effects of the energy dissipation mechanism are felt slowly, a solution of the same form may be assumed but with $w_1 = w_1(t)$, a slowly varying function of time. Substituting this new solution into equations (16) and (17), and averaging over the period $t = 2\pi/\sigma_1 \sqrt{(\sigma_1/\sigma_1)}$ eliminates all but the secular terms and results in.

$$w_{1}\dot{w}_{0} = -2\sigma_{2}h_{0}(I_{p}\dot{w}_{p} + I_{r}\dot{w}_{r})/(I_{1}\dot{f}\sigma_{1} + I_{2}\dot{f}\sigma_{2})$$
(13)
$$\dot{F} = w_{0}\dot{w}_{0}(I_{1}\sigma_{1} + I_{2}\sigma_{2})/2\sigma_{1} + I_{p}w_{p}\dot{w}_{p} + I_{r}w_{r}\dot{w}_{r} + I_{r}w_{r}\dot{w}_{r} + I_{r}w_{r}\dot{w}_{p}$$
(19)

Substituting equation (18) into equation (19) and introducing,

$$\sigma_{0} = h_{0}(I_{1}\sigma_{1} + I_{2}\sigma_{2})/(I_{1}^{2}\sigma_{1} + I_{2}^{2}\sigma_{2})$$
(20)

with a little manipulation produces,

$$\dot{T} = -(I_{p} - I_{r})\dot{w}_{p}(\sigma_{0} - w_{r}) + - I_{r}(\dot{w}_{p} + \dot{w}_{r})(\sigma_{0} - (w_{p} + w_{r}))$$
(21)
$$= P_{p} + P_{r}$$

where P_p and P_r are the platform and rotor components of the energy dissipation rate. It is necessary that,

 $P_p < 0$ $P_r < 0$

Now, let

$$\sigma_{\rm p} = \sigma_{\rm p} - W_{\rm p} \tag{22}$$

$$\sigma_r = \sigma_0 - (w_r + w_r) \tag{23}$$

so that,

$$P_{z} = -(I_{y} - I_{z})\dot{w}_{y}\sigma_{z} \qquad (24)$$

$$P_{1} = -I_{1}(\dot{w}_{1} + \dot{w}_{1})\sigma_{1}$$
 (25)

or alternately.

$$P_{p}/\sigma_{p} = -(I_{p} - I_{r})\dot{w}_{p}$$
(26)

$$P_r/\sigma_r = -I_r(\dot{w}_p + \dot{w}_r)$$
 (27)

Substituting equations (26) and (27) into equation (18) yields,

$$w_{0}\dot{w}_{0} = \left[2\sigma_{2}h_{0} / (I_{1}^{2}\sigma_{1} + I_{2}^{2}\sigma_{2}) \right] \left[(P_{p} / \sigma_{p}) + (P_{r} / \sigma_{r}) \right]$$
(28)

As a necessary and sufficient condition for stability,

$$w_{n}\dot{w}_{n} < 0 \tag{29}$$

Since $\sigma_1 \sigma_2 > 0$ is also necessary for stability and $h_1 > 0$ by convention, it follows that,

$$(P_{p}/\sigma_{p}) + (P_{r}/\sigma_{r}) < 0$$
 (30)

And because P_p and P_r are both negative, at least one of the two of σ_p and σ_r must be positive such that the total effective energy dissipation rate of equation (30) is negative.

In terms of the specific application to +his study the equations simplify as follows. For a system with a despun platform and no damping mechanism on the platform, w_p is

essentially zero and P_{p} is zero. Therefore,

$$\sigma_1 = I_r w_r / I_1 \tag{3}$$

$$\sigma_2 = I_{\rm r} w_{\rm r} / I_2 \tag{9}$$

and,

$$\sigma_{1} = I_{w}(I_{1}\sigma_{1} + I_{2}\sigma_{2})/(I_{1}\sigma_{1} + I_{2}\sigma_{2})$$
(20)

$$h_{0} = I_{v}w_{v} \tag{15}$$

Now,

$$\sigma_{p} = \sigma_{p} \tag{22}$$

$$\sigma_{\mu} = \sigma_{\mu} - w_{\mu} \qquad (23)$$

and the stability criteria is,

$$\mathbf{P}_{t}/\sigma_{t} < 0 \tag{30}$$

Or, since P, must be negative,

$$\sigma_{\star} > 0 \tag{31}$$

Substituting equations (8) and (9) into equation (20) yields,

$$\sigma_0 = 2I_r w_r / (I_1 + I_1)$$
 (32)

and substituting equation (32) into equation (23) gives,

$$\sigma_{r} = 2I_{r}w_{r}/(I_{1} + I_{2}) - w_{r}$$
(33)

or, from equation (31),

$$w_r[2I_r/(I_1 + I_2) - 1] > 0$$
 (34)

Dividing by w_{r} and adding one to both sides of equation (34) yields,

$$2I_{r}/(I_{1} + I_{2}) > 1$$
 (35)

which is the familiar spin axis moment of inertia divided by the algebraic mean transverse moment of inertia developed by Likins. It is at this point that Spencer asserts that using the geometric mean in the denominator yields more accurate results.

Spencer's conclusion is based on the fact that w^3 , the rotor-fixed nutation frequency, actually varies over time as a function of I_1 , the transverse moment of inertia. And he states that I: varies as,

$$I_{1}(\hat{w}t) = I_{1}\sin^{2}\hat{w}t + I_{1}\cos^{2}\hat{w}t \qquad (36)$$

where.

$$\hat{w} = [I_{1}/((I_{1}I_{2}))]w_{1}$$
 (37)

from equations (10) and (11). These variations in w^3 must be reflected in \hat{T} . Likins neglected this in averaging over the period to obtain equations (18) and (19). Unfortunately, this makes the solution for an asymmetric satellite much more complicated. Spencer does not provide a complete solution. Instead, he presents a simplified example and shows that using the geometric mean for I_{\uparrow} provides a closer approximation to simulation results.

B. SIMULATION DESCRIPTION

The computer simulation used in this study to determine the numerical solution to the equations of motion of the satellite was developed as a part of the work done by Chung in Reference 1. This section provides a description of this simulation. A dual spin spacecraft can be modelled as two rigid bodies capable of rotating relative to each other about a common axis. This common axis passes through the mass center of the body representing the rotor. The liquid in spherical fuel tanks can be modelled as an axisymmetric spherical pendulum with the hinge point at the center of the tank. The spherical pendula can be mounted on either of the two bodies that constitute the system (Figure 2). The energy dissipation due to the liquid sloshing can be included in the model by introducing viscous damping in the spherical joint of the pendula.



Figure 2. Schematic of a generic dual spin spacecraft.

Chung applied Kane's method to the model described above to determine the equations governing the dynamic behavior of the system. This method entails defining generalized forces called inertia forces and active forces. These forces are expressed as functions of the generalized speeds (rotational velocities) of the various components of the system. The total force acting on the system can be summarized as,

 $F_{r} + F_{r}^{\star} = 0$ (r = 1, 2, 3,...,N) (38)

where N is the total number of generalized speeds, and \mathbb{F}_{r} and \mathbb{F}_{r}^{t} are the total generalized active and inertia forces in inertial space. The total number of generalized speeds, N, is equal to the number of degrees of freedom of the system. In this case, three for the three components of the angular velocity of the platform, plus one for the relative velocity between the platform and the rotor, plus three times the number of fuel tanks (each pendulum representing a fuel tank has three degrees of freedom).

To obtain the generalized inertia forces of the system, the contributions from the platform, rotor, and pendula are summed. To obtain the generalized active forces of the system, the contributions from all active forces on each part of the system are summed. It is assumed that the resultant of the external forces of the platform and the rotor are zero

and that no external forces act on the pendula mounted on either the platform or the rotor.

In addition to the generalized speeds obtained by solving the differential equations developed using Kane's method, several other quantities are useful in understanding the motion of the spacecraft. These include the central angular momentum of the system, the kinetic energy of the system, the energy dissipated through the spherical joints of the pendula, the work done by the motor and the external forces, and the nutation angle of the system. All of these can be expressed in terms of the generalized speeds.

The simulation takes as input the system parameters that characterize the properties of the satellite being simulated. These include the mass, moments of inertia, location and orientation for each component of the system relative to the center of mass, as well as the key properties of the pendula (length and damping coefficient). The initial conditions for all of the generalized speeds and coordinates must also be provided. The output of the simulation is a set of values that characterize the state of the system at a given point in time. These include the nutation angle, the kinetic energy, the total energy, the work done by the rotor and the external forces, the two components of the transverse angular velocity, the platform angular velocity and the rotor angular velocity.

C. SATELLITE CONFIGURATION

The configuration of the satellite used in this study is a derivative of the INTELSAT VI satellite. The satellite consists of a platform and a symmetric rotor. There are four fuel tanks and four oxidizer tanks mounted on the platform in the configuration shown in Figure 3.

.



Figure 3. Arrangement of fuel and oxidizer tanks.

The liquid characteristics are as follows.

fuel density = 376.2 kg/m²

oxidizer density = 1448.3 kg/m^2

fuel kinematic viscosity = 9.73x10⁷

oxidizer kinematic viscosity = 2.92x10⁻⁷

The mass and inertia properties of the INTELSAT VI satellite [Ref. 1] are shown in Table I.

	PLATFORM (DRY)	ROTOR (DRY)			
	,		25.2	20	
MASS(Kg)	1058.9	695.7	2503.1	2326.0	2193.1
I _{zz} (Kg-m ²)	1587.1	927.0	4469.5	4182.5	1950.E
Im (Kg-m ²)	1518.3	1166.0	4491.2	4098.7	3118.3
I ₁₁ (Xg-m ²)	1529.4	973.7	4458.5	4171.5	3939.E
In (Kg-m ²)	0.0	0.0	0.0	0.0	0.0
$I_{22}(Kg-m^2)$	44.4	-6.1	38.3	39.3	38.3
$I_{XY}(Kg-m^2)$	0.0	0.0	0.0	0.0	0.0

TABLE I. INTELSAT VI mass and inertia properties.

These parameters are modified using the procedure described in Chapter III, Section A to obtain the configuration used in the study (Tables II-IV).

	PLATFORM (DRY)	ROTOR (DRY)	TOTAL
MASS(Kg)	1058.2	695.7	2503.1
I _{XX} (Kg-m ²)	1095.2	1418.9	4469.5
I _{yy} (Kg-m ²)	1518.3	2791.0	6116.2
I(Kg-m ⁻)	1084.2	1413.9	4459.5
I	0.0	0.0	Û Û
(<u>Ka-m</u> -)	44.4	-6.1	د قر
I _m (Kg-m ²)	0.0	9.0	0.0

TABLE II. Study satellite mass and inertia properties - - 26.2% fuel load.

TABLE III. Study satellite mass and inertia properties - -20% fuel load.

	PLATFORM (DRY)	ROTOR (DRY)	TOTAL
MASS(Kg)	1058.9	695.7	2326.0
$I_{\pi}(Kg-m^2)$	1006.0	1467.3	4192.5
$I_{yy}(Kg-m^2)$	1518.3	2387.9	5820.6
I ₁₁ (Kg-m ²)	1035.8	1467.3	4171.5
$I_{\gamma z}(Kg-m^2)$	0.0	0.0	2.0
I _{XZ} (Kg-m ²)	44.4	-6.1	38.3
$I_{xy}(Kg-m^2)$	0.0	0.0	0.0

	PLATFORM (DRY)	ROTOR (DRY)	TOTAL
MASS(Kg)	1058.8	695.7	2133.1
I _{XX} (Kg-m ²)	1005.0	1508.1	3950.6
I(Xg-m ²)	1518.3	2969.4	5581.7
I(Kg-m ²)	995.0	1508.1	3939.6
In (Fg-m ²)	0.0		
I. (Yg-m ¹)	44.4	- 6. 1	33.3
I. (Kg-m ⁻)	٥.٥	2.0	0.0

TABLE IV. Study satellite mass and inertia properties - - 15% fuel load.

III. PROCEDURE

This chapter describes the methodology used over the course of the study. The first section deals with how to determine the spacecraft confirmation and system parameters. Section B describes how the energy sink criteria is applied to a specific configuration to determine the stability of that configuration. Section C contains the procedure for maning the simulation and obtaining the simulation output.

A. SYSTEM PARAMETERS

In order to investigate the behavior of a marginally stable spacecraft it is first necessary to model a spacecraft with this configuration. To accomplish this, the system parameters of the INTELSAT VI satellite were modified to obtain the appropriate configuration. One of the programs written by Chung in Reference 1 to support the simulation produces as its output a summary of the INTELSAT VI system parameters. To modify the system parameters to achieve a given inertia ratio, I_g/\tilde{I}_1 , where $\tilde{I}_1 = (I_1 + I_2)/2$, the algebraic mean of the spacecraft transverse moments of inertia ratio to determine the desired axial moment of inertia, I_s .

moment of inertia of the wet rotor, I... However, the simulation takes as its input the dry spacecraft parameters. Subtracting the old wet rotor moment of inertia from the desired moment of inertia yields the amount the dry rotor axial moment of inertia must be increased. It is a general theorem of rigid body mechanics that for a given body, the sum of any two of the principle moments of inertia must be greater than the third. This must be kept in mind to achieve a realistic design. To adhere to this principle, the sum of the dry rotor transverse moments of inertia is subtracted from the new dry rator axial moment of inertia. The result is divided by two and added to each of the dry rotor transverse moments of inertia. An additional adjustment is made to achieve perfect symmetry on the rotor. The difference between the dry rotor transverse moments of inertia is added to the smaller of the two to make them equal. To balance the system and maintain the same total spacecraft transverse moments of inertia, the corresponding platform transverse moments of inertia are decremented by the same amounts as were added to the rotor transverse moments of inertia.

B. ENERGY SINK PREDICTION

Given numerical values for I_1 , I_2 , I_p , I_r , w_p , and w_r , developing an energy sink prediction is a reasonably straight forward application of the equations derived in Chapter II.

The best way to implement the energy sink equations is by means of a simple spread sheet using the equations below.

$$\mathbf{h}_{2} = \mathbf{I}_{p}\mathbf{w}_{p} + \mathbf{I}_{p}\mathbf{w}_{p}$$
(15)

$$\sigma_{1} = [(I_{p} - I_{1})w_{p} + I_{r}w_{r}]/I_{1}$$
(3)

$$\sigma_{1} = [(I_{p} - I_{1})w_{p} + I_{r}w_{r}]/I_{1}$$
(9)

$$\sigma_{1} = h_{1}(I_{1}\sigma_{1} + I_{1}\sigma_{1})/(I_{1}\sigma_{1} + I_{1}\sigma_{1})$$
(20)

$$\sigma_{1} = \sigma_{1} - w_{1} \qquad (22)$$

$$\sigma_{1} = \sigma_{1} \sim (w_{1} + w_{1}) \qquad (23)$$

Recall that the stability criteria were,

$$\mathfrak{s},\mathfrak{s},\mathfrak{r} > 0 \tag{12}$$

$$(P_{p}/\sigma_{p}) + (P_{p}/\sigma_{p}) < 0$$
 (30)

and that P₁ and P₁ are always negative. Concentrating on the second relationship, for the system being studied, P₁ is zero and P₁ is negative since all of the fuel tanks are mounted on the rotor. Therefore, if σ_1 is positive the system is inherently stable. However, if σ_1 is negative σ_2 must be positive and a nutation damper must be installed on the platform to overcome the destabilizing effect of the rotor.

C. SIMULATION PROCEDURE

To determine the stability of the system the parameter of interest is the nutation angle, Φ . If Φ grows without bound, the system is unstable. If Φ damps out to zero, the system is stable. In executing the simulation, an initial transverse

rate of 0.2 radians per second is applied to induce an initial nutation angle.

The first step in the study entails finding a satellite configuration with an inertia ratio slightly greater than one for which the simulation results indicate the system is stable. An inertia ratio of 1.01 is arbitrarily chosen for the initial run. If the simulation results indicate the system is unstable the inertia ratio must be increased and the simulation run again. This process is continued until a stable configuration is found upon which the remainder of the study will be based. After letermining a stable configuration the platform asymmetry is varied and the satellite motion is simulated for three different fuel loads. The INTEISAT VI satellite uses a liquid propellant rocket for its apogee kick motor and consumes nearly 75% of its fuel during this maneuver. The beginning of life fuel load for the INTELSAT VI satellite upon which the study configuration is based is 26.2%. Therefore, fuel loads of 26.2%, 20%, and 15% were chosen for this study. The initial asymmetry was arbitrarily selected to be 5%. If the results indicate the system is stable the asymmetry is increased by 5% and the simulation run again. This process continues until the simulation results indicate the system is unstable.

IV. RESULTS AND ANALYSIS

This chapter presents the results of the study and a discussion of their significance. The first section presents a summary of the results of the simulation runs. Section B contains a summary of the energy sink predictions for all of the cases simulated.

A. SIMULATION RESULTS

As outlined in chapter III, the first step in the study was to find a configuration for which the system was stable. The first simulation run was for the case of a symmetric system with an inertia ratio of 1.01 and a fuel load of 26.2%. The simulation time was set at 200 seconds. The result is chown in Figure 4.



Figure 4. Initial simulation result.

From the graph it is difficult to determine whether the system is stable or unstable because the simulation time was not of sufficient length. Rather than extending the simulation time to arrive at a more conclusive result, the simulation was run again with the damping coefficients increased by a factor of 100. Figure 5 shows that this time the results are much more definitive.



Figure 5. Initial simulation result with increased damping.

The system is obviously unstable. From this point on, all simulation runs were made with the damping coefficients increased in order to minimize the main frame CPU time.

Next the simulation was run with an inertia ratio of 1.1. The result reveals that the system is stable as shown in Figure 6.



Figure 6. Result for an inertia ratio of 1.1 and a fuel load of 26.2%.

Having achieved stability, the fuel load was changed to 20% and the simulation was run again. The system remained stable as shown in Figure 7 so the fuel load was reduced to 15% and the simulation run again. The result still indicated that the system was stable (Figure 8). Having achieved stability for all three fuel loads, the next step was to begin varying the asymmetry of the platform.



Figure 7. Result for an inertia ratio of 1.1 and a fuel load of 20%.



Figure 8. Result for an inertia ratio of 1.1 and a fuel load of 15%.

A summary of the results for all of the asymmetry variations can be seen in Table V. The system remained stable for every case simulated.

FUEL LOAD	5%	10%	ASYM! 15%	METRY 20%	25%	35%	55%
15%	s	s	S	S	s	s	S
20%	S	S	S	S	S	S	S
26.2%	S	S	S	S	S	S	S
	Not	:e: U	= Uns	table	e, S =	= Stał	ole

TABLE V. Simulation results, $2I_s/(I_1 + I_2) = 1.1$.

In light of the results reported in Reference 7, the results in Table V were surprising. However, upon reexamining the energy sink equations, it was discovered this is exactly what should have been expected. Starting with a symmetric satellite, such that $I_1 = I_2 = I_1$ the stability criteria is,

$$2I_{r}/(I_{1} + I_{2}) > 1$$
 (35)

To create an asymmetry, let $I_1 = I_t + a$ and $I_2 = I_t - a$. Substituting for I_1 and I_2 in equation (35) reduces to the identical equation indicating the system should remain stable. Using the geometric mean as suggested by Spencer, equation (35) becomes,

$$I_{\mu}/([I_{\mu}I_{2}) > 1)$$
 (29)

Substituting for I- and I- here yields,

$$I_{*}/(\sqrt{(I_{*}^{2} - a^{2})}) > 1$$
 (40)

which has to be greater than the original inertia ratio indicating that the system is getting more stable as the asymmetry is increased. Regardless of the method used to determine the average transverse moment of inertia, both predict the system should remain stable because $I_{\rm F}/\bar{I}_{\rm c}$ is greater than one. This being the case, it was now necessary to find a configuration which was unstable to determine if the system would become stable as the asymmetry was increased.

The simulation was run for symmetric configurations with inertia ratios of 1.03, 1.05, and 1.07. The results are summarized in Table VI. They show that for a symmetric spacecraft the stability cutoff falls between inertia ratios of 1.03 and 1.05.

LOAD	1.01	1.03	1.05	1.07	1.1
 15%		U	S	S	s
20%	-	U	S	S	S
26.2%	υ	U	S	S	s

TABLE VI. Simulation results of varying inertia ratio.

Given the results in Table VI, the configuration with an inertia ratio of 1.03 was selected to evaluate the influence of varying the asymmetry.

The summary of these results is provided in Table VII. As expected, the system eventually became stable as the asymmetry was increased.

FUEL		ASYM	AETRY	
LOAD	0%	25%	40%	55%
58	IJ	IJ	S	S
20%	IJ	11	S	S
. 6 . 2%	U	U	S	S

TABLE VII. Simulation results for varying the platform asymmetry.

What appeared to have happened was that as the asymmetry was increased, the inertia ratio increased enough to become greater than the stability cutoff inertia ratio as equation (40) indicates. To confirm this, a graph of inertia ratio versus asymmetry was created using equation (39) to employ the geometric mean transverse moment of inertia (Figure 9). As expected, it showed that for a symmetric inertia ratio of 1.03, as the platform asymmetry was increased, the inertia ratio increased enough to become greater than the stability
cutoff inertia ratio that falls between 1.03 and 1.05. This plot combined with Table VII indicate that the stability cutoff inertia ratio falls between 1.032 and 1.035.



Figure 9. Inertia ratio vs. asymmetry for 26.2%, 20%, and 15% fuel loads.

To try and pin down the stability cutoff, the simulation was run for the symmetric case, for all three fuel loads with inertia ratios of 1.035, 1.04 and 1.045. Adding the results to Table VI shows that the stability cutoff inertia ratio for a symmetric satellite is different than for an asymmetric satellite. The simulation results indicate the stability cutoff for the symmetric satellite is approximately 1.045 (Table VIII).

FUEL LOAD	1.01 1		INE 1.035	RTIA F 1.04	RATIO 1.045	1.05	1.07	1.10
15%		U	U	U	M	S	S	S
20%	-	U	U	U	м	S	S	S
25.2%	IJ	U	U	U	М	S	S	S
N	ote: V =	Unst	able,	s = st	able,	M = Ma	rginal	

TABLE VIII. Simulation results for varying inertia ratio of symmetric spacecraft.

Finally, several simulation runs were made with fuel loads of 50% and 75% to investigate the effect of higher fuel loads. Higher fuel loads make the system more stable. In fact these higher fuel loads lowered the stability cutoff inertia ratio for a symmetric satellite to approximately 1.02 (Table IX).

TABLE IX. Simulation results of varying inertia ratio for higher fuel loads.

FUEL LOAI	1.01	INE 1.02	RTIA RA' 1.03	FIO 1.05	1.07	1.1
50%	U	M	S	S	S	S
75%	U	M	S	S	S	S
	Note: U =	Unstable,	S = Sta	able, M =	= Margin	al

The results of varying the asymmetry on a configuration with an inertia ratio of 1.01 and a fuel load of 75% are shown in Table X.

FUEL LOAD	0%	ASYMMETRY 25%	40%	
75%	U	U	S	
	Note; U = Unst	able, S = Stable	2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	

TABLE X. Simulation results of varying asymmetry for a 75% fuel load.

Here, as with the lower fuel loads the system became stable as the asymmetry was increased. Graphing inertia ratio versus asymmetry for a configuration with a 75% fuel load shows that the stability cutoff inertia ratio in this case is between 1.012 and 1.015 (Figure 10). As with the symmetric cases, the stability cutoff inertia ratio is significantly lower for the higher fuel loads than for the lower fuel loads.





B. ENERGY SINK PREDICTIONS

As was shown in Chapter II, the energy sink stability criteria is,

$$w_{n}\dot{w}_{n} < 0 \tag{29}$$

and,

$$(P_p/\sigma_p) + (P_r/\sigma_r) < 0 \tag{30}$$

where one or the other, or both, of σ_p or σ_r must be positive such that equation (30) is true. For the specific case of a dual spin spacecraft with a despun platform, equation (30) reduces to.

 $I_{r}/\tilde{I}_{r} > 1$

By definition, all of the energy sink predictions for this study were stable since the initial requirement in defining the configuration was an inertia ratio slightly greater than one. It should be noted that it is irrelevant which method is used to calculate the average transverse moment of inertia. The energy sink predictions were incorrect for every case where the simulation results showed the system was unstable. For symmetric configurations with the lower fuel loads, the energy sink predictions were incorrect when the inertia ratio was less than 1.045. For symmetric configurations with the higher fuel loads, the energy sink predictions were incorrect when the inertia ratio was less than 1.02. For the asymmetric variations on a configuration with a symmetric inertia ratio

of 1.03. with the lower fuel loads. the energy sink predictions were incorrect for asymmetries less than 25%, correct for asymmetries greater than 40%, and indeterminate for the transition zone between 25% and 40% asymmetry. Using Spencer's method to determine the inertia ratio, 40% asymmetry equates to an inertia ratio of 1.035. 25% asymmetry equates to an inertia ratio of 1.032.

V. SUMMARY AND CONCLUSIONS

This chapter presents the conclusions based on the results presented in the Chapter IV.

A. SUMMARY

The energy sink stability criteria for a dual spin spacecraft with a despun platform specify that the system must have an inertia ratio greater than one. The results of this study indicate that it is not sufficient for the system's inertia ratio just to be greater than one. It must be greater than one by a predictable amount. Below this stability cutoff inertia ratio the energy sink method predicts stability when the simulation results indicate the system is unstable. Further, the stability cutoff inertia ratio varies inversely with the fuel load. For the case of the lower fuel loads used in this study the stability cutoff inertia ratio is approximately 1.045 for a symmetric spacecraft. For the higher fuel loads it is approximately 1.02 (Figure 11).

TIVERTIA PATIO VO TODE LOAD



Figure 11. Inertia ratio vs. fuel load showing for stability, inertia ratio must be greater than one.

As shown in the development in Chapter II. Likins' stability criteria can be reduced to show that the key parameter for stability predictions is the ratio of the spin moment of inertia to the algebraic mean transverse moment of inertia. Recall that Spencer plaimed that Likins was incorrect and that the geometric mean transverse moment of inertia is the key stability parameter. The results of this study support Spencer, indicating that as the asymmetry is increased the system becomes more stable. This trend is predictable using the geometric mean in computing the inertia ratio. Interestingly, the stability cutoff inertia ratio for

the asymmetric variations on a satellite with a symmetric inertia ratio of 1.03 is around 1.035. As the asymmetry is increased, the inertia ratio grows from the symmetric 1.03 until it eventually reaches 1.05 for a 75% platform asymmetry. It exceeds the stability cutoff inertia ratio at 40% asymmetric (Figure 12). It must be pointed out that while Spencer's geometric mean transverse moment of inertia more accurately shows the effects of asymmetry, it is also inaccurate in predicting stability below the stability cutoff inertia ratio.



Figure 12. Inertia ratio vs. asymmetry. Calculating inertia ratio using Likins' method vs. Spencer's method.

B. CONCLUSIONS

It is unclear why there is a difference in the stability cutoff for the symmetric and asymmetric configurations. This is perhaps an opportunity for future research. What has been determined or confirmed in this study is that:

(1) the energy sink stability criteria is not valid below a stability cutoff inertia ratio which is not just equal to one, but greater than one.

(2) stability increases as the fuel load is increased.

(3) stability increases as the platform asymmetry is increased.

(4) Spencer was correct in asserting that it is more accurate to use the geometric mean transverse moment of ipertia than the algebraic mean transverse moment of ipertia in computing the inertia ratio.

APPENDIX A

SIMULATION DATA

Included here are the graphs of nutation angle versus time for each simulated case. The graphs are ordered in the sequence in which they are presented in the text.

















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SYM, 15/1t = 1 03, 15% FUEL

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25% ASYM, Is/It = 1.03, 26.2% FUEL

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40% ASYM, IS/It = 1.03, 26 2% FUEL



40% ASYM, Is/It = 1.03, 15% FUEL



57M, 16/11 = 1.035, 20% FUEL









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SYM, 1s/1t = 1.03, 50% FUEL















25% ASYM, Is/It = 1 03, 75% FUEL













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APPENDIX B

ENERGY SINK PREDICTIONS

Included in this appendix are the spread sheet data for the energy sink predictions. The spread sheet takes as its input, w_p , w_r , I_1 , I_2 , I_p , and I_r . The spread sheet uses the following equations to calculate σ_r ,

$$h_0 = I_p w_p + I_r w_r \tag{15}$$

$$\sigma_1 = [(I_p - I_2)w_p + I_rw_r]/I_1$$
(8)

$$\sigma_2 = [(I_p - I_1)w_p + I_rw_r]/I_2$$
(9)

$$\sigma_0 = h_0 (I_1 \sigma_1 + I_2 \sigma_2) / (I_1^2 \sigma_1 + I_2^2 \sigma_2)$$
(20)

 $\sigma_{\rm p} = \sigma_0 - w_{\rm p} \tag{22}$

$$\sigma_{\rm r} = \sigma_0 - (w_{\rm p} + w_{\rm r}) \tag{23}$$

For the configurations used in this study, as long as σ_r is positive, the energy sink criteria predicts stability.

Sigmal Sigma2 Sigma0 SigmaP SigmaR Predict Result Wr II I2 EUN WO, WP Ip Ir ho (22) (23) (15) (9) (20) EQUATION MUMPER (8) SYMMETRIC, Is/10 = 1.0, 75% FUEL 20030601 0.000072 3.141519 6499.4 6488.4 8012.2 6493.9 20401.29 3.138878 3.144199 3.141609 3.141536 0.000017 STABLE UNSTABLE SYMMETRIC, Is/It = 1.0, 50% FUEL 90020602 0 000072 3.141517 5795.1 5984.1 7507.9 5989.6 18816.99 3.138656 3.144425 3.141611 3.141538 0.000018 STABLE UNSTABLE

 M'N NO.
 No
 Nr
 II
 IZ
 Ip
 Ir
 Sigmal
 Sigmal

SYMMETRIC, Is/It = 1.005, 75% FUEL P0020902 0.000072 3.141519 6409.4 6408.4 9044.7 6526.4 00503 40 3.154567 3.159935 3.157332 3.157259 0.015739 STABLE UNSTABLE SYMMETRIC, Is/It = 1.005, 50% FUEL P0020901 0.000072 3.141519 5995.1 5984.1 7537.8 6019.5 18910.92 3.154324 3.160122 3.157293 3.157221 0.015701 STABLE UNSTABLE

SYMMETRIC, Is/It = 1.01, 75% FUEL 90031301 0.0000/2 3.141519 6499.4 6489.4 9077.1 6558.8 20605.18 3.170249 3.175623 3.173006 3.172933 0.031414 STABLE UNSTABLE SYMMETRIC, Is/It = 1.01, 50% FUEL 90031302 0.000072 3.141519 5995.1 5984.1 7557.8 6049.5 19005.17 3.170045 3.175872 3.173029 3.172956 0.031436 STABLE UNSTABLE SYMMETRIC, Is/It = 1.01, 26.2% FUEL 90021501 0.000072 3.141519 4469.5 4458.5 6026.9 4508.6 14164.29 3.169028 3.176846 3.173005 3.172932 0.031412 STABLE UNSTABLE SYMMETRIC, Is/It = 1.01, 20% FUEL

0.000072 3.141519 4182.5 4171.5 5737.1 4218.8 13253.86 3.168612 3.177168 3.173057 3.172984 0.031464 STABLE SYMMETRIC, Is/It = 1.01, 15% FUEL

0.000072 3.141519 3950.6 3939.6 5502.8 3984.5 12517.78 3.168506 3.177352 3 172995 3.172923 0.031403 STABLE 25% ASYMMETRIC, Is/It = 1.01, 75% FUEL 90031301 0.000072 3.141519 6872.3 6115.5 8077.1 6558.8 20605.18 2.998231 3.369256 3.173006 3.172933 0.031413 STABLE UNSTABLE 40% ASYMMETRIC, Is/It = 1.01, 75% FUEL 90031302 0.000072 3.141519 7096.1 5891.7 8077.1 6558.8 20605.18 2.903673 3.497237 3.173006 3.172933 0.031413 STABLE STABLE

SUM NO.WpWrI1I2IpIrboSigma1Sigma2Sigma0SigmaPSigmaRPredictResultEQUATION NUMBER(15)(8)(9)(20)(22)(23)SYMMETRIC, Is/It = 1.1, 75% FUEL

0.000072 3.141519 5499.4 6488.4 8661.6 7143.3 22441.44 3.452776 3.458630 3.455773 3.455701 0.314181 STABLE STABLE SYMMETRIC, Is/It = 1.1, 50% FUEL

0.000072 3.141519 5995.1 5984.1 8106.9 6588.6 20698.80 3.452548 3.458894 3.455791 3.455718 0.314198 STABLE STABLE SYMMETRIC, Is/It = 1.1, 26.2% FUEL 90022001 0.000072 3.141519 4469.5 4458.5 6428.7 4910.4 15426.58 3.451451 3.459966 3.455776 3.455703 0.314183 STABLE STABLE STABLE SYMMETRIC, Is/It = 1.1, 201 FUEL 20022003 0.000072 3.141519 4182.5 4171.5 6113 4594.7 14424.78 3.451161 3.460261 3.455778 3.455705 0.314185 STABLE STABLE STABLE SYMMETOTO, Ta/IL = 1.1. 15% FIEL 70022002 0.00072 3.141519 3950.6 3939.6 5857.9 4339.6 13633.26 3.450888 3.460523 3.455771 3.455699 0.314179 STABLE STABLE 5% ASYMMETRIC, Is/It = 1.1, 26.2% FUEL 90022004 0.000072 3.141519 4516.2 4411.8 6428.7 4910.4 15426.59 3.415752 3.496590 3.455776 3.455703 0.314183 STABLE STABLE 5% ASYMMETRIC, Is/It = 1.1, 20% FUEL 90022005 0.000072 3.141519 4227.3 4126.7 6113 4594.7 14434.78 3.414587 3.497826 3.455778 3.455705 0.314185 STABLE STABLE 5% ASYMMETRIC, Is/It = 1.1, 15% FUEL 90022101 0.000072 3.141519 3933.7 3836.5 5957.9 4339.6 13633.36 3.413647 3.498800 3.455771 3.455699 0.314179 STABLE STABLE 10% ASYMMETRIC, Is/It = 1.1, 26.2% FUEL 90022102 0.000072 3.141519 4562.9 4365.1 6428.7 4910.4 15426.58 3.380803 3.533998 3.455776 3.455703 0.314183 STABLE STABLE 10% ASYMMETRIC, Is/It = 1.1. 20% FUEL 90022103 0.000072 3.141519 4272 4082 6113 4594.7 14431.78 3.378859 3.536128 3.455778 3.455705 0.314185 STABLE STABLE 10% ASYMMETRIC, IS/It = 1.1. 15% FUEL 90022104 0.000072 3.141519 4036.8 3853.4 5857.9 4339.6 12633.36 3.377201 3.537933 3.455771 3 455699 0.314179 STABLE STABLE 15% ASYMMETRIC, Is/It = 1.1, 26.2% FUEL

90022105 0.000072 3.141519 4609.5 4318.5 6428.7 4910.4 15426.58 3.346626 3.572131 3.455776 3.455703 0.314183 STABLE STABLE

Wr 11 12 Ip Ir ho Sigma1 Sigma2 Sigma0 SigmaP SigmaR Predict Result RUN NO. WO (15) (8) (3) (20) (22) (23) EQUATION NUMBER 15% ASYMMETRIC, Is/It = 1.1, 20% FUEL 90022105 0.000072 3.141519 4316.8 4037.2 6113 4594.7 14434.79 3.343794 3.575367 3.455778 3.455705 0.314185 STABLE STABLE 15% ASYMMETRIC, Is/It = 1.1, 15% FUEL 90022107 0.000072 3.141519 4080 3810.2 5857.9 4339.6 13633.36 3.341443 3.578045 3.455771 3.455699 0.314179 STABLE STABLE 201 ASYMMETRIC, IS/It = 1.1, 26.21 FUEL 90022108 0.000072 3.141519 4656.2 4271.8 6428.7 4910.4 15426.58 3.313061 3.611182 3.455776 3.455703 0.314183 STABLE STABLE 20% ASYMMETRIC. Is/It = 1.1, 20% FUEL 90022109 0.00072 3.141519 4361.5 3992.5 6113 4594.7 14434.78 3.309525 3.615336 3.455778 3.455705 0.314185 STABLE STABLE 20% ASYMMETRIC, Is/It = 1.1, 15% FUEL 90022110 0.000072 3.141519 4123.1 3767.1 5957.9 4339.6 13633.36 3.306514 3.618981 3.455771 3.455699 0.314179 STABLE STABLE 25% ASYMMETRIC, Is/It = 1.1, 26.2% FUEL 20022111 0.000072 3.141519 4702.9 4225.1 6428.7 4910.4 15426.58 3.280163 3.651095 3.455776 3.455703 0.314193 STABLE STABLE 25% ASYMMETRIC, Is/It = 1.1, 20% FUEL 90022112 0.000072 3.141519 4406.3 3947.7 6113 4594.7 14434.78 3.275877 3.656424 3.455778 3.455705 0.314185 STABLE STABLE 25% ASYMMETRIC. Is/It = 1.1. 15% FUEL 90022113 0.000072 3.141519 4166.2 3724 5857.9 4339.6 13633.36 3.272309 3.660865 3.455771 3.455698 0.314179 STAPLE STAPLE 35% ASYMMETRIC, Is/It = 1.1, 26.2% FUEL 90022201 0.000072 3.141519 4850.9 4077.1 6428.7 4910.4 15426.58 3.180088 3.783629 3.455776 3.455703 0.314183 STABLE STABLE 35% ASYMMETRIC, Is/It = 1.1, 20% FUEL 90022202 0.000072 3.141519 4547 3807 6113 4594.7 14434.78 3.174512 3.791556 3.455777 3.455705 0.314185 STABLE STABLE 35% ASYMMETRIC, Is/It = 1.1, 15% FUEL 90022203 0.000072 3.141519 4300.8 3589.4 5857.9 4339.6 13633.36 3.169899 3.798142 3.455771 3.455698 0.314178 STABLE STABLE 55% ASYMMETRIC, Is/It = 1.1, 26.2% FUEL 90022204 0.000072 3.141519 4796.2 4131.8 6428.7 4910.4 15426.58 3.216355 3.733539 3.455776 3.455703 0.314183 STABLE STABLE

Sigmal Sigma2 Sigma0 SigmaP SigmaR Predict Result RUN NO SP Hr II I2 Ip Ir ho EQUATION NUMBER (15)(8) (?) (20) (22) (23) 55% ASYMMETRIC, IS/It = 1.1, 20% FUEL 20022201 0.000072 3.141519 4495.8 3858.2 6113 4594.7 14434.78 3.210664 3.741241 3.455777 3.455705 0.314185 STABLE STABLE 55: ASYMMETRIC, IS/It = 1.1, 15: FUEL 20022202 0.000072 3.141519 4252.5 3637.7 5857.9 4339.6 13633.36 3.205902 3.747713 3.455771 3.455698 0.314178 STABLE STABLE SYMMETRIC, Is/It = 1.03, 75% FUEL 90022701 0.000072 3.141519 6499.4 6488.4 8207 6688.7 21013.28 3.233038 3.238519 3.235849 3.235776 0.094256 STABLE STABLE SYMMETRIC, Is/It = 1.03, 50% FUEL 90022601 0.000072 3.141519 5995.1 5984.1 7687.6 6169.3 19381.53 3.232823 3.238766 3.235865 3.235792 0.094272 STABLE STABLE SYMMETRIC. Is/It = 1.03, 26.2% FUEL 90022402 0.000072 3.141519 4469.5 4458.5 6116.2 4597.9 14444.83 3.231796 3.239769 3.235851 3.235778 0.094258 STABLE UNSTABLE SYMMETRIC, Is/It = 1.03, 20% FUEL 90022401 0.00072 3.141519 4182.5 4171.5 5820.6 4302.3 13516.18 3.231531 3.240052 3.235859 3.235786 0.094266 STABLE UNSTABLE SYMMETRIC, IS/IL = 1.03, 15% FUEL 90022304 0.000072 3.141519 3950.6 3939.6 5581.7 4063.4 12755.65 3.231248 3.240270 3.235826 3.235753 0.094233 STABLE UNSTABLE 25% ASYMMETRIC, Is/It = 1.03, 75% FUEL 70022901 0.000072 3.141519 6856.1 6131.7 8207 6688.7 21013.28 3.064837 3.426909 3.235848 3.235776 0.094256 STABLE STABLE 25% ASYMMETRIC, IS/It = 1.03, 50% FUEL 90022902 0.000072 3.141519 6296.2 5683 7687.6 6169.3 19381.53 3.078225 3.410360 3.235864 3.235792 0.094272 STABLE STABLE 25% ASYMMETRIC, Is/It = 1.03, 26.2% FUEL 90022501 0.000072 3.141519 4741.9 4186.1 6116.2 4597.9 14444.83 3.046149 3.450585 3.235850 3.235778 0.094258 STABLE UNSTABLE 25% ASYMMETRIC, IS/IL = 1.03, 20% FUEL 90022502 0.000072 3.141519 4442.8 3911.2 5820.6 4302.3 13516.18 3.042203 3.455681 3.235859 3.235786 0.094266 STABLE UNSTABLE 25% ASYMMETRIC, ID/IL = 1.03, 15% FUEL 90022503 0.000072 3.141519 4200.7 3689.5 5581.7 4063.4 12765.65 3.038872 3.459913 3.235825 3.235753 0.094233 STABLE UNSTABLE

RUN NO. WP Wr -11 12 Ip Ir ho Sigmal Sigma2 Sigma0 SigmaP SigmaR Predict Result EQUATION NUMBER (8) (9) (20) 115 (22) (23) 40: ASYMMETRIC, IS/It = 1.03, 75% FUEL 90022804 0.000072 3.141519 7070.1 5917.7 9207 6688.7 21013.28 2.972072 3.550833 3.235848 3.235775 0.094255 STABLE STABLE 40% ASYMMETRIC, Is/It = 1.03, 50% FUEL 90030101 0.000072 3.141519 6476.9 5502.3 7687.6 6169.3 19281.53 2.992347 3.522357 3.235864 3.235791 0.094272 STAPLE STAPLE 40% ASYMMETRIC, Is/It = 1.03, 26.2% FUEL 20030102 0.000072 3.141519 4905.4 4022.6 6116.2 4597.9 14444.83 2.944621 3.590832 3.235850 3.235777 0.094257 STABLE STABLE 40% ASYMMETRIC, Is/It = 1.03, 20% FUEL 90030103 0.000072 3.141519 4599 3755 5820.6 4302.3 13516.18 2.938880 3.593427 3.235858 3.235785 0.034265 STABLE STABLE 40% ASYMMETRIC, Is/It = 1.03, 15% FUEL 90030104 0.000072 3.141519 4350.8 3539.4 5591.7 4063.4 12765.65 2.934035 3.606639 3.235825 3.235752 0.094232 STAPLE STAPLE 55% ASYMMETRIC, Is/It = 1.03, 75% FUEL 0.000072 3.141519 7284.1 5703.7 8207 6688.7 21013.28 2.884758 3.684056 3.235848 3.235775 0.034255 STABLE 55% ASYMMETRIC, Is/It = 1.03, 50% FUEL

90022803 0.000072 3.141519 6657.6 5321.6 7687.6 6169.3 17381.53 2.911131 3.641959 3.235864 3.235791 0.094271 STABLE STABLE 55% ASYNHETRIC, Is/It = 1.03, 26.2% FUEL •

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90022504 0.000072 3.141519 5068.9 3859.1 6116.2 4597.9 14444.83 2.849643 3.742963 3.235849 3.235777 0.094257 STABLE STABLE STABLE 55% ASYMMETRIC, Is/It = 1.03, 20% FUEL

90022505 0.000072 3.141519 1755.2 3598.8 5820.6 4302.3 13516.18 2.842345 3.755651 3.235857 3.235785 0.094265 STABLE STABLE 55% ASYMMETRIC, Is/It = 1.03, 15% FUEL

90022506 0.000072 3.141519 4500.9 3389.3 5581.7 4063.4 12765.65 2.836190 3.766361 3.235824 3.235752 0.094232 STABLE STABLE STABLE

SYMMETRIC, Is/It = 1.035. 26.2% FUEL 90030603 0.000072 3.141519 4469.5 4458.5 6138.5 4620.2 14514.89 3.247471 3.255483 3.251544 3.251472 0.109952 STABLE UNSTABLE SYMMETRIC, Is/It = 1.035, 20% FUEL

90030604 0.000072 3.141519 4182.5 4171.5 5841.5 4323.2 13581.84 3.247230 3.255792 3.251578 3.251505 0.109985 STABLE UNSTABLE

RUN NO. Wp Wr II I2 Ip Ir ho Sigmal Sigma2 SigmaO SigmaP SigmaR Predict Result EQUATION NUMBER (15) (8) (7) (20) (22) (23) SYMMETRIC, Is/It = 1.035, 15% FUEL 90030605 0.000072 3.141519 3950.6 3939.6 5601.5 4083.2 12827.86 3.246994 3.256060 3.251593 3.251520 0.110000 STABLE UNSTABLE SYMMETRIC, Is/It = 1.04, 25.2% FUEL 90030701 0.000972 3.141519 4469.5 4458.5 6160.9 4642.6 14595.26 3.263216 3.271266 3.267309 3.267236 0.125716 STABLE UNSTABLE SYMMETRIC, Is/It = 1.04. 20% FUEL 90030702 0.000072 3.141519 4182.5 4171.5 5862.4 4344.1 13647.50 3.262928 3.271532 3.267297 3.267297 3.267295 0.125705 STABLE UNSTABLE SYMMETRIC, Is/It = 1.04, 15; FUEL 90030703 0.000072 3.141519 3950.6 3939.6 5621.2 4102.9 12889.75 3.262659 3.271769 3.267208 0.125688 STABLE UNSTABLE SYMMETRIC, Is/It = 1.045, 26.2% FUEL 90020903 0.000072 3.141519 4469.5 4458.5 6183.2 4664.9 14655.32 3.278890 3.286980 3.283003 3.282930 0.141410 STABLE MARGINAL SYMMETRIC, Is/It = 1.045, 20% FUEL 90030804 0.000072 3.141519 4182.5 4171.5 5883.3 4365 13713.16 3.278627 3.287272 3.283017 3.282944 0.141424 STABLE MARGINAL SYMMETRIC, Is/It = 1.045, 15% FUEL 90030905 0.000072 3.141519 3950.6 3939.6 5640.9 4122.6 12951.64 3.278325 3.287479 3.282968 3.282896 0.141376 STABLE MARGINAL SYMMETRIC, Is/It = 1.05, 75% FUEL 90022702 0.000072 3.141519 6499.4 6488.4 8336.9 6819.6 21421.37 3.295827 3.301415 3.298618 0.157099 STAPLE STAPLE SYMMETRIC, Is/It = 1.05, 50% FUEL 9:0022602 0.000072 3.141519 5995.1 5984.1 7807.4 6289.1 19757.90 3.295602 3.301660 3.298701 3.298628 0.157108 STABLE STABLE SYMPETRIC, Is/It = 1.05, 26.2% FUEL 90022308 0.000072 3.141519 4469.5 4458.5 6205.5 4687.2 14725.38 3.294565 3.302693 3.298697 3.298624 0.157104 STARLE STARLE SYMPETRIC, Is/It = 1.05, 20% FUEL 90022309 0.000072 3.141519 4182.5 4171.5 5904.1 4385.8 13778.50 3.294250 3.302937 3.298661 3.299588 0.157068 STABLE STABLE

ho Sigmal Sigma2 Sigma0 SigmaP SigmaR Predict Result end no Mb II I2 Ip Ir Nr (15) CONVERSION NUMBER (7) (20) (22) (23)(8) SYMMETRIC, 10/11 = 1.05, 153 FUEL 20022203 0.000072 3.141519 3950.6 3939.6 5660.7 4142.4 13013.84 3.294071 3.203268 3.298736 3.299653 0.157143 STARLE STARLE 25% ASYMMETRIC, IS/It = 1.05, 25.2% FUEL 90022405 0.000072 3.141519 4730.8 4197.2 6205.5 4687.2 14725.38 3.112597 3.508300 3.298696 3.298624 0.157104 STAPLE STAPLE 25% ASYMMETRIC, Is/It = 1.05, 20% FUEL 20022406 0.000072 3.141519 1432.4 3921.6 5904.1 4385.8 13778.50 3.108524 2.513409 3.229560 3.239558 0.157068 STAPLE STAPLE 25% ASYMMETRIC, IS/It = 1.05, 15% FUEL 20022407 0 000072 3.141519 4120.9 2699.3 5660.7 4142.4 13013.94 3.105198 3.517838 3.298725 3.298663 0.157143 STABLE STAPLE SYMMETRIC, Is/It = 1.07, 75% FUEL 20022703 0.000072 3.141519 6499.4 6488.4 8466.8 6948.5 21829.46 3.358616 3.364310 3.361534 3.361461 0.219941 STABLE STABLE SYMMETRIC, Is/It = 1.07, 50% FUEL 20022603 0.000072 3.141519 5795.1 5794.1 7727.2 6408.9 20134.26 3.358380 3.364553 3.361537 3.361464 0.217944 STABLE STABLE SYMMETRIC, Is/It = 1.07, 26.2% FUEL 90222403 0.000072 3.141519 4469.5 4458.5 6294.8 4775.5 15005.92 3.357333 3.365616 3.361542 3.361470 0.219950 STABLE STABLE SYMMETRIC. IS/IL = 1.07. 20% FUEL 20022404 0.000072 3.141519 4182.5 4171.5 5987.7 4469.4 14041.14 3.357045 3.365897 3.361538 3.361465 0.219945 STABLE STABLE SYMEIRIC, Is/It = 1.07, 15% FUEL 90022305 0.000072 3.141519 3950.6 3939.6 5739.6 4221.3 13261.71 3.356813 3.366186 3.361566 3.361493 0.219973 STAPLE STAPLE SYMMETRIC, IS/It = 1.02, 75% FUEL 00012703 0.000072 3.141519 6499.4 6488.4 8142.1 6623.8 20809.39 3.201667 3.207095 3.204452 3.204379 0.062859 STABLE STABLE STABLE SYMMETRIC, Is/It = 1.02, 50% FUEL 9C022403 0.000072 3.141519 5995.1 5984.1 7627.7 6109.4 19193.35 3.201434 3.207319 3.204447 3.204374 0.062854 STABLE STABLE

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LIST OF REFERENCES

- Stanford University, Final Report Project No. INTEL 465, Energy sink Analysis of Dual Spin Spacecraft, by Ting-Mong Chung, October 1985.
- Likins. Peter W., "Attitude Stability Criteria for Dual Spin Spacecraft," Journal of Spacecraft and Rocket, v. 4, pp. 1638-1643, December 1967.
- Thomson, W.T. and Reiter, G. S., "Motion of an Asymmetric Body with Internal Dissipation," AIAA Journal. v. 1. no. 6. pp. 1429-1430, June 1963.
- Spender, T. M., "Energy-Sink Analysis for Asymmetric Dual-Spin Spacecraft," Journal of Spacecraft and Rocket, v. 11, pp. 463-468, July 1974.
- 5. Hubert, Carl, "Spacecraft Attitude Acquisition from an Arbitrary Spinning or Tumbling State," Journal of Guidance and Control, v. 4, pp. 164-170, March-April 1981.
- Cochran, J. E. Jr. and Shu, P. H., "Effects of Energy Addition and Dissipation on Dual-Spin Spacecraft," Journal of Guidance and Control, v. 6, pp. 368-373, September-October 1983.
- 7. International Telecommunications Satellite Organization (INTELSAT), Study of INTELSAT VI During Transfer and On-Orbits, Due to Asymmetries in Platform and Rotor, by Sudeep K. Singh, 3 November 1986.

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